

# Energy balance model

## Building models

- What is the actual problem?
- How much complexity?  
*(i.e., define scope)*
- What are the state variables?  
*(i.e., define system)*
- Are there conservation laws?
- Are there rules for changes in state variables?  
*(i.e., what is the mathematical model)*
- How should this be implemented in a computer?  
*(i.e., discrimination, numerical methods)*
- What are useful diagnostics?  
*(i.e., output that is needed to answer science objective)*
- How do you test the model is correct?
- How do you double check the model is correct?

# Global energy balance model

- i.e., radiative equilibrium

$$c_p \frac{dT}{dt} = (1 - \alpha) \frac{S}{4} - \varepsilon \sigma T^4$$

Allows us to predict global mean temperature if we know the albedo, emissivity and the solar constant

## Time stepping

- Can write change as:  
change in state over some time  $\frac{dT}{dt} \approx \frac{T_{\text{future}} - T_{\text{now}}}{\text{time difference}} = \frac{T_{\text{future}} - T_{\text{now}}}{\Delta t}$
- That is, a **finite difference**
- As such, we can make a prediction for one (small) time step ( $\Delta t$ )
- The advance in time (future becomes now) by one time step  $T_{\text{future}} = T_{\text{now}} + \Delta t \frac{dT}{dt}$
- and start again...

*This is an integration of the differential equation in time.*

## Today's objectives

1. How to log onto atoc
2. How to write and compile a Fortran program
3. How to write and compiler a useful Fortran program
4. Think about some experiments with your shiny new EBM

*Remember use class web site and our wiki as a first source of help.*

# Experiments/discussion

- What is the radiative equilibrium temperature of Earth?
- Does this depend on the choice of the time step?
- Change the heat capacity so as to model the upper 70 meters of the ocean rather than the atmosphere ( $C_p = 2.95 \times 10^8 \text{ J/K}$ ). Does the equilibrium temperature change?
- Is the mean temperature the same if there is a diurnal cycle? Annual cycle? "Glacial" cycle?
- If the sun should stop, how long does it take for the temperature to become a factor of  $e^{-1}$  its equilibrium value (i.e., what is the e-folding time), for both the atmosphere and upper ocean?
- How can the e-folding time be found analytically from (1)?
- What are some limitations of this model, and what are their consequences?
- How would you construct a coupled atmosphere-ocean model with this level of complexity?

For next Tuesday's class